

# Advances of the ASTRI MINI-ARRAY - the CTA Precursor & the Small Size Telescopes (SSTs) of CTA



Elisabete M. de Gouveia Dal Pino and Maria Victoria del Valle  
(On behalf of the CTA and ASTRI Mini-Array Group in Brazil)





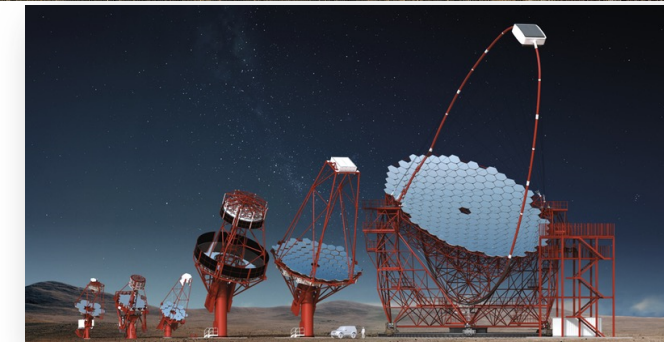
Universidade de São Paulo

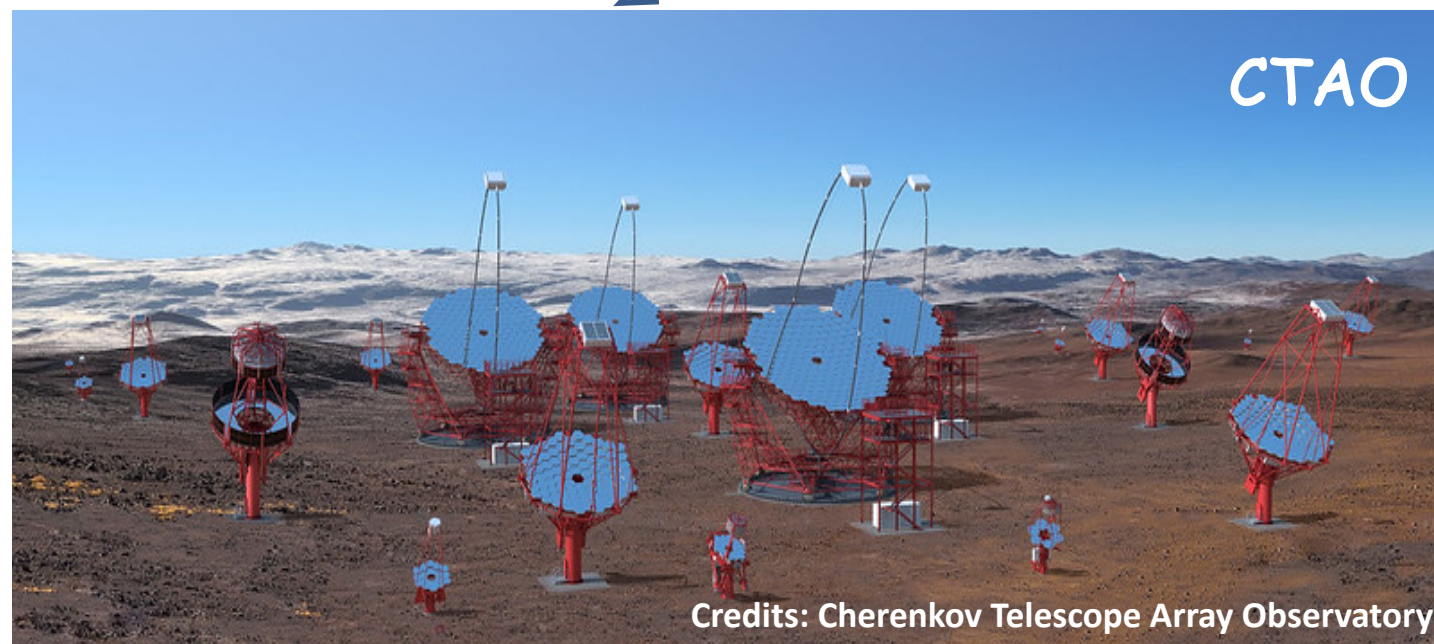


## ASTRI Mini-Array with 9 Cherenkov telescopes:

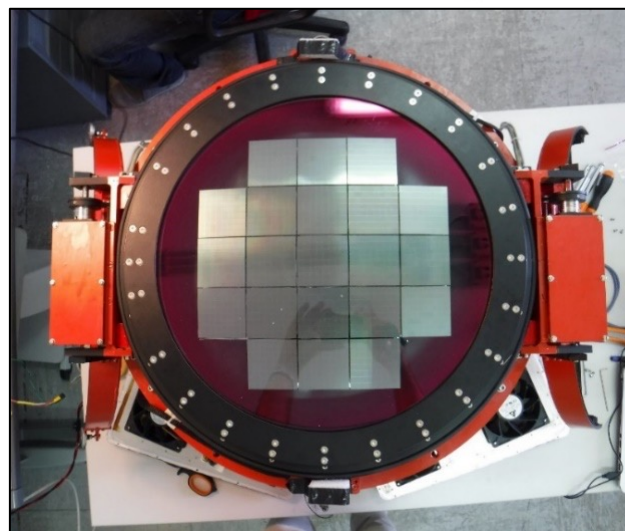
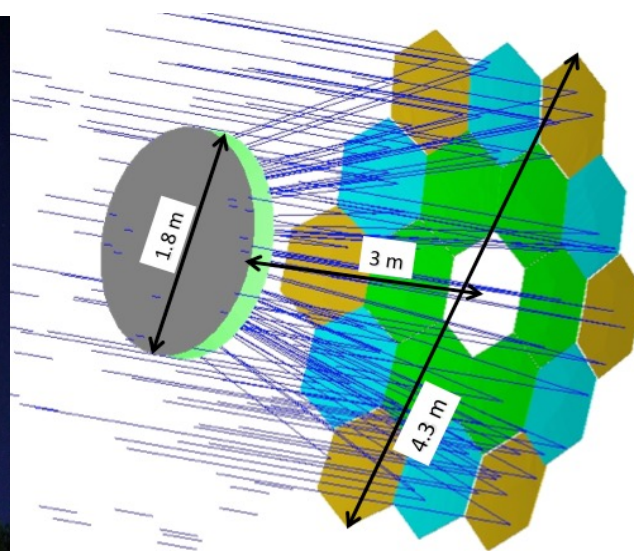
for ground based high-energy gamma-ray astronomy up to few 100 TeV

- **Partners: Italy, Brazil, South Africa, Spain**
  - INAF, INFN, and several Italian Universities - more than 100 hundred researchers
  - **Universidade de São Paulo/FAPESP (Brazil): provided 3 of the telescopes (mechanical structures)**
  - North-West University (South Africa)
  - Recently: IAC (Spain) for site support in Tenerife





1. Development of an end-to-end technological **PROTOTYPE** for CTA
2. Construction and operation of the **ASTRI mini-array @** Teide Observatory (Canary Islands)
3. Leading the consortium that will **implement the SST telescopes @ CTAO** southern site



- **Opto-mechanics**
  - Alto-azimuthal mount
  - **Schwarzschild-Couder optical design**
  - Primary Mirror: 4.3 m (segmented)
  - Secondary Mirror: 1.8 m (monolithic)
  - F/#: 0.5 m
  - Average effective area: 5.0 m<sup>2</sup>
  - Optical PSF  $\leq 0.19$  deg
  - Post calibration pointing precision  $\leq 7$  arcsec
- **Cherenkov Camera**
  - **Front End electronics based on CITIROC ASICs**
  - **SiPM sensors: 7x7 mm**
  - 1344 pixels (2368 when focal plane fully populated)
  - Field of View: 8.2 deg (10.5 deg)
  - Angular pixel size: 0.19 deg
- **Expected performance**
  - Energy threshold  $\approx 1$  TeV
  - Energy/Angular resolution  $< \sim 25\%$  /  $< \sim 0.15^\circ$
  - Sensitivity  $\approx 1$  Crab @  $5 \sigma$  in few hours



**24<sup>th</sup> September 2014**

**Inauguration of the prototype @ INAF-Catania mountain station in Serra La Nave placed at 1725 meters on the Etna volcano**



**25<sup>th</sup> of May 2017**

## First Cherenkov light with the ASTRI camera

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**cta** cherenkov telescope array

Press Release

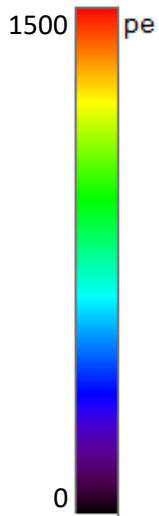
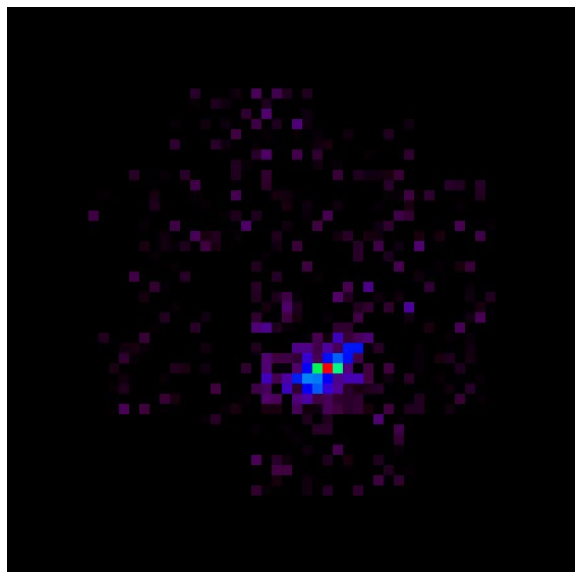
### CTA Prototype Telescope, ASTRI, Achieves First Light

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During the nights of 25 and 26 May, the camera of the ASTRI telescope prototype (pictured to the left) recorded its first ever Cherenkov light while undergoing testing at the astronomical site of Serra La Nave (Mount Etna) in Sicily managed by INAF-Catania. This comes not long after its optical validation was achieved in November 2016 ([read story here](#)). This accomplishment was the first optical demonstration for astronomical telescopes using the novel Schwarzschild Couder dual-mirror design. The ASTRI telescope is a proposed Small-Sized Telescope design for the Cherenkov Telescope Array (CTA).

Although the camera was not fully configured, the ASTRI team was still able to capture its first Cherenkov light and produce beautiful images of the showers generated by cosmic rays in the Earth's atmosphere. The image below shows one of the events captured by the team. This information will allow scientists to reconstruct the direction of gamma-ray photons emitted from celestial sources (indicated by the yellow line on the image on the left). The camera is based on novel SiPM small pixel sensors (7 mm x 7 mm) and CITIROC ASICS peak-finder front-end electronics. The camera was specifically designed to fit on the dual mirror ASTRI telescopes for covering a large field of view of 10° x 10°.



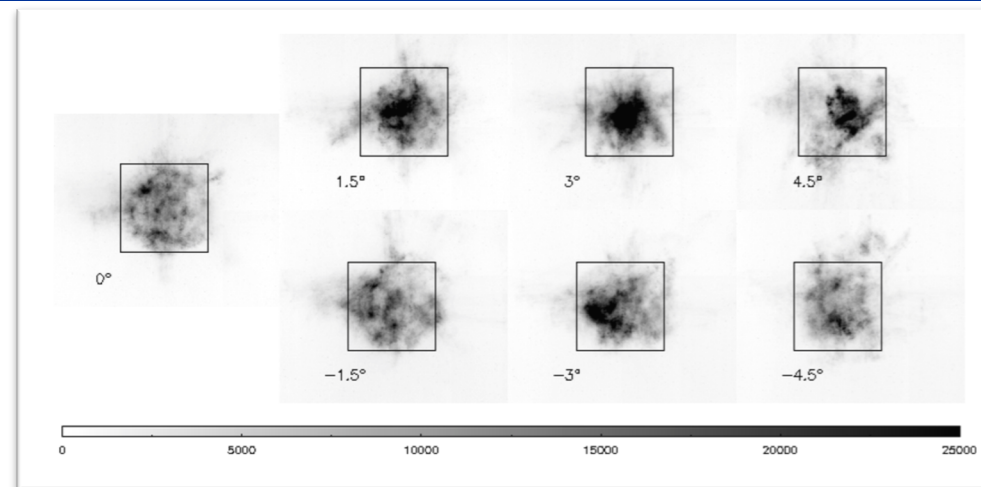


A&A 608, A86 (2017)  
DOI: 10.1051/0004-6361/201731602  
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Astronomy  
&  
Astrophysics

## First optical validation of a Schwarzschild Couder telescope: the ASTRI SST-2M Cherenkov telescope

E. Giro<sup>1,2</sup>, R. Canestrari<sup>2</sup>, G. Sironi<sup>2</sup>, E. Antolini<sup>3</sup>, P. Conconi<sup>2</sup>, C. E. Fermino<sup>4</sup>, C. Gargano<sup>5</sup>, G. Rodeghiero<sup>1,6</sup>,  
F. Russo<sup>7</sup>, S. Scuderi<sup>8</sup>, G. Tosti<sup>3</sup>, V. Vassiliev<sup>9</sup>, and G. Pareschi<sup>2</sup>

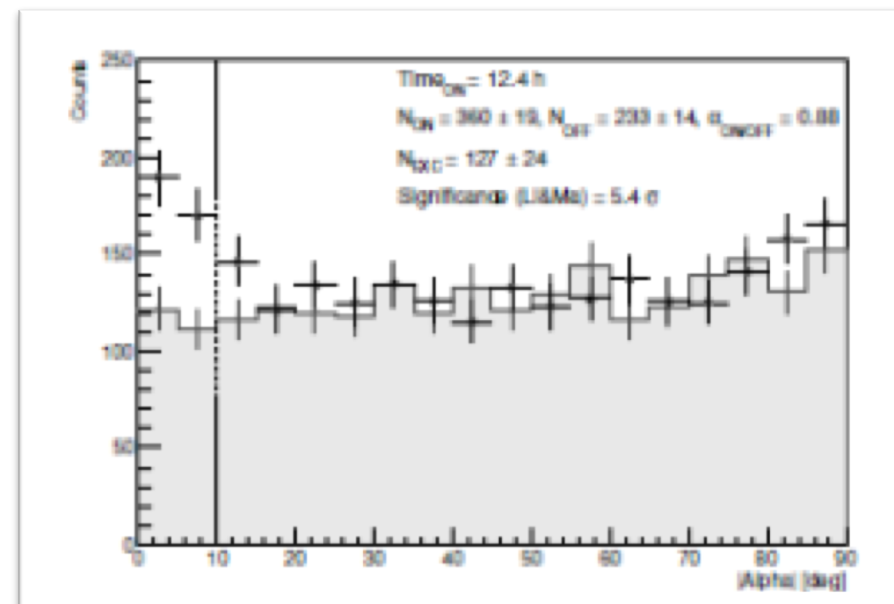


A&A 634, A22 (2020)  
<https://doi.org/10.1051/0004-6361/201936791>  
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Astronomy  
&  
Astrophysics

## First detection of the Crab Nebula at TeV energies with a Cherenkov telescope in a dual-mirror Schwarzschild-Couder configuration: the ASTRI-Horn telescope

S. Lombardi<sup>1,2,\*</sup>, O. Catalano<sup>3,\*</sup>, S. Scuderi<sup>4,\*</sup>, L. A. Antonelli<sup>1,2</sup>, G. Pareschi<sup>5</sup>, E. Antolini<sup>6</sup>, L. Arrabito<sup>7</sup>,  
G. Bellasai<sup>8</sup>, K. Bernlöhr<sup>9</sup>, C. Bigongiari<sup>1</sup>, B. Biondo<sup>3</sup>, G. Bonanno<sup>8</sup>, G. Bonnoli<sup>5</sup>, G. M. Böttcher<sup>10</sup>, J. Bregeon<sup>11</sup>,  
P. Bruno<sup>8</sup>, R. Canestrari<sup>3</sup>, M. Capalbi<sup>3</sup>, P. Caraveo<sup>4</sup>, P. Conconi<sup>5</sup>, V. Conforti<sup>12</sup>, G. Contino<sup>3</sup>, G. Cusumano<sup>3</sup>,  
E. M. de Gouveia Dal Pino<sup>13</sup>, A. Distefano<sup>4</sup>, G. Farisato<sup>14</sup>, C. Fermino<sup>13</sup>, M. Fiorini<sup>4</sup>, A. Frigo<sup>14</sup>, S. Gallozzi<sup>1</sup>,  
C. Gargano<sup>3</sup>, S. Garozzo<sup>8</sup>, F. Gianotti<sup>12</sup>, S. Giarrusso<sup>3</sup>, R. Gimenes<sup>13</sup>, E. Giro<sup>14</sup>, A. Grillo<sup>8</sup>, D. Impiombato<sup>3</sup>,  
S. Incorvaia<sup>4</sup>, N. La Palombara<sup>4</sup>, V. La Parola<sup>3</sup>, G. La Rosa<sup>3</sup>, G. Leto<sup>8</sup>, F. Lucarelli<sup>1,2</sup>, M. C. Maccarone<sup>3</sup>,  
D. Marano<sup>8</sup>, E. Martinetti<sup>8</sup>, A. Miccichè<sup>8</sup>, R. Millul<sup>5</sup>, T. Mineo<sup>3</sup>, G. Nicotra<sup>15</sup>, G. Occhipinti<sup>8</sup>, I. Pagano<sup>8</sup>,  
M. Perri<sup>1,2</sup>, G. Romeo<sup>8</sup>, F. Russo<sup>3</sup>, F. Russo<sup>12</sup>, B. Sacco<sup>3</sup>, P. Sangiorgi<sup>3</sup>, F. G. Saturni<sup>1</sup>, A. Segreto<sup>3</sup>, G. Sironi<sup>5</sup>,  
G. Sottile<sup>3</sup>, A. Stamerra<sup>1</sup>, L. Stringhetti<sup>4</sup>, G. Tagliaferri<sup>5</sup>, M. Tavani<sup>16</sup>, V. Testa<sup>1</sup>, M. C. Timpanaro<sup>8</sup>, G. Toso<sup>4</sup>,  
G. Tosti<sup>17</sup>, M. Trifoglio<sup>12</sup>, G. Umam<sup>8</sup>, S. Vercellone<sup>5</sup>, R. Zanmar Sanchez<sup>8</sup>, C. Arcaro<sup>14</sup>, A. Bulgarelli<sup>12</sup>,  
M. Cardillo<sup>16</sup>, E. Cascone<sup>18</sup>, A. Costa<sup>8</sup>, A. D'Ai<sup>3</sup>, F. D'Ammando<sup>12</sup>, M. Del Santo<sup>3</sup>, V. Fioretti<sup>12</sup>, A. Lamastra<sup>1</sup>,  
S. Mereghetti<sup>4</sup>, F. Pintore<sup>4</sup>, G. Rodeghiero<sup>14</sup>, P. Romano<sup>5</sup>, I. Schwarz<sup>5</sup>, F. Sciaccia<sup>8</sup>, F. R. Vitello<sup>8</sup> and A. Walter<sup>5</sup>



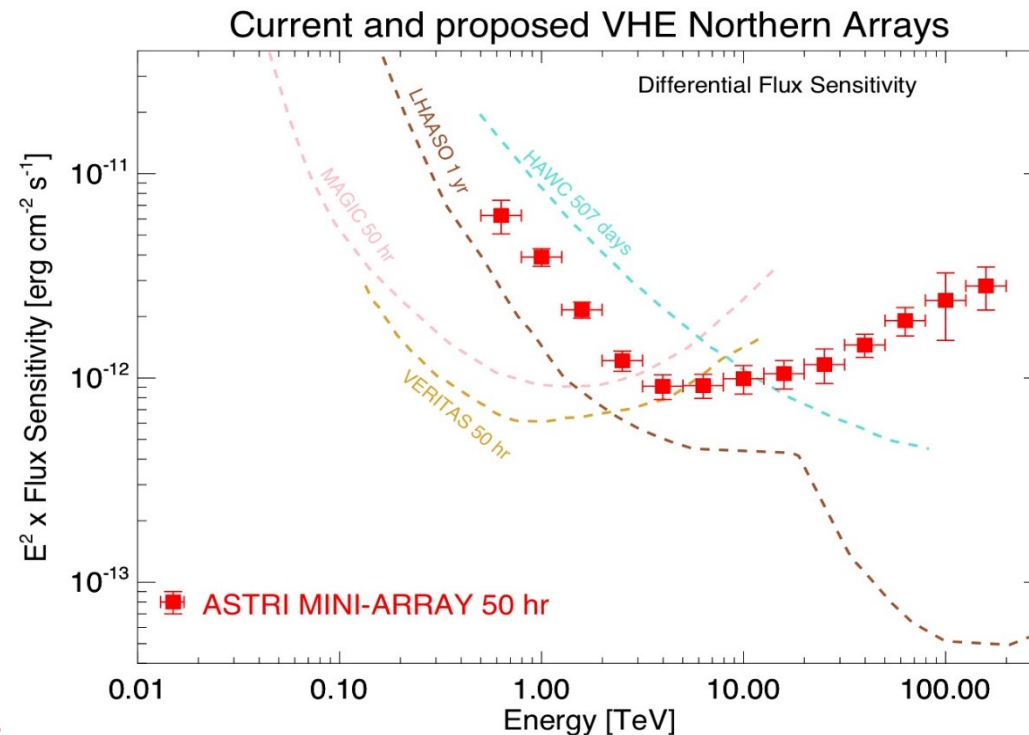


**The ASTRI mini-array will be a new pathfinder of the arrays of Cherenkov telescopes**

- **INAF commitment with the Italian government and international partners: **Brazil**, South Africa, Spain**
- **Dedicated funding**
- **It will have 9 ASTRI telescopes: resulting from **evolution** of the ASTRI-Horn prototype successfully implemented and tested**
- **It will be deployed at the Teide Observatory (Canary Islands) in collaboration **with Spain & Brazil****

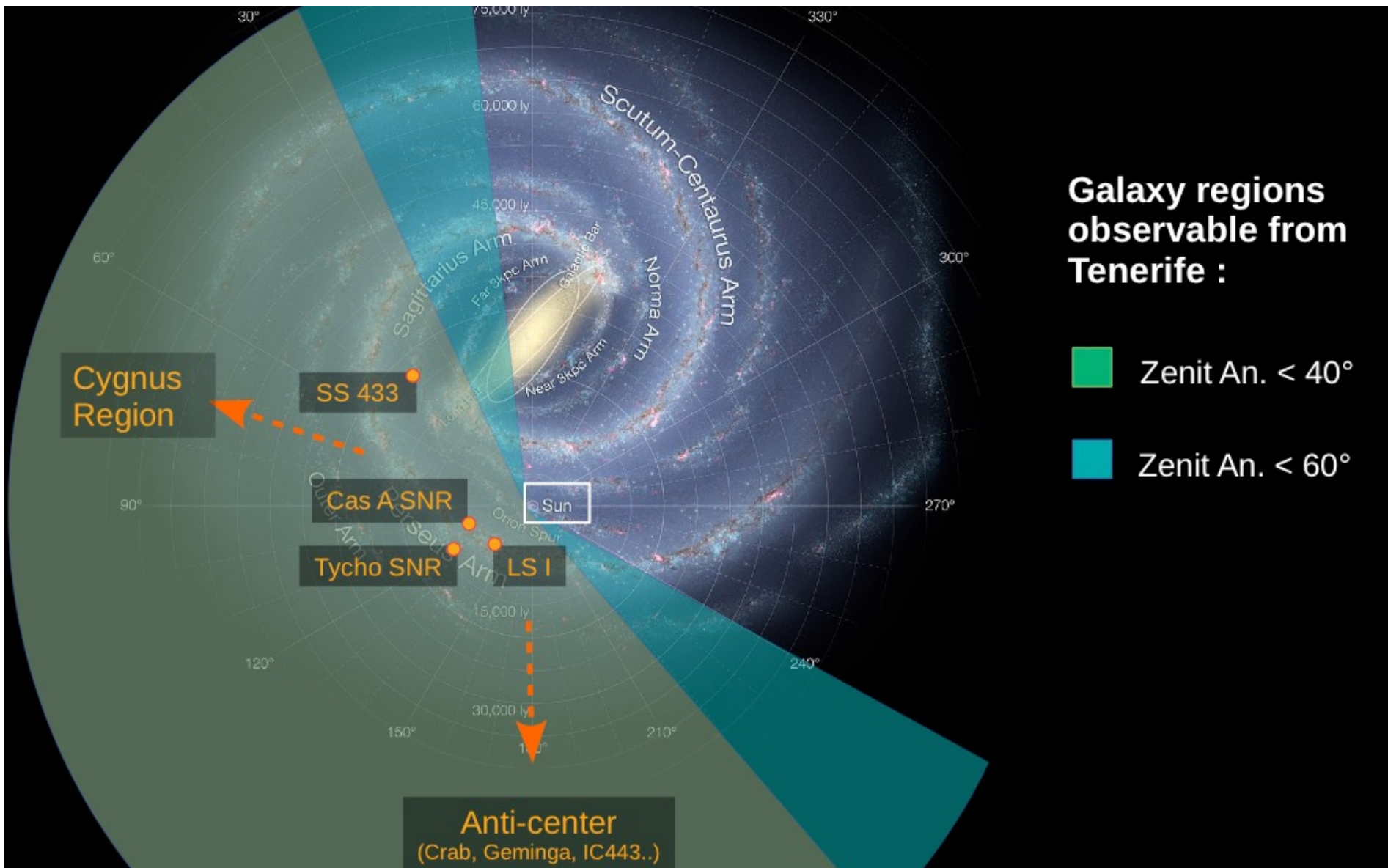


## Science at very high gamma ray energy in the northern hemisphere!



### Expected performance:

- **Sensitivity: better than current IACTs ( $E > 10$  TeV):**
  - Possibility to extend the spectra of already detected sources and/or measure cut-offs
  - Possibility to characterize the morphology of extended sources at the highest VHE
- **Energy/Angular resolution:  $< \sim 10\%$  /  $< \sim 0.1^\circ$  ( $E > \sim 10$  TeV)**
- **Wide FoV ( $\geq 10^\circ$ ), with homogeneous off-axis acceptance**
  - Optimal for multi-target fields, surveys, and extended sources
  - Enhanced chance for serendipity discoveries

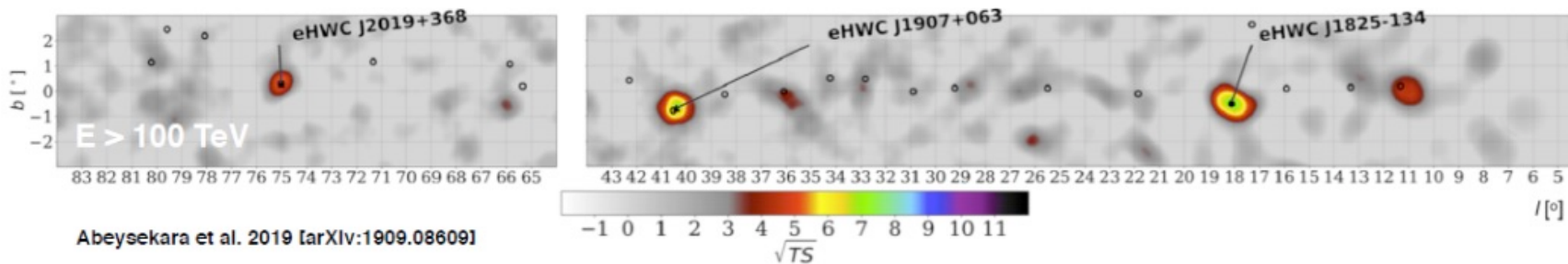
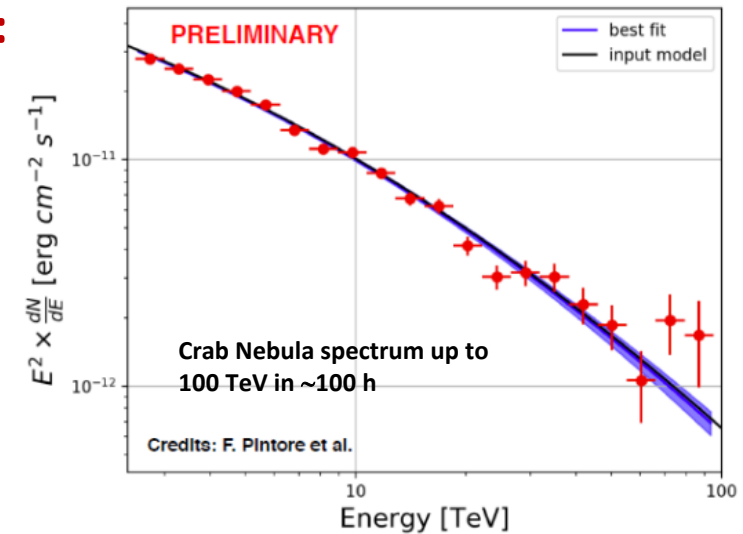


## Core science program in the first ~2/3 years (detailed strategy under definition):

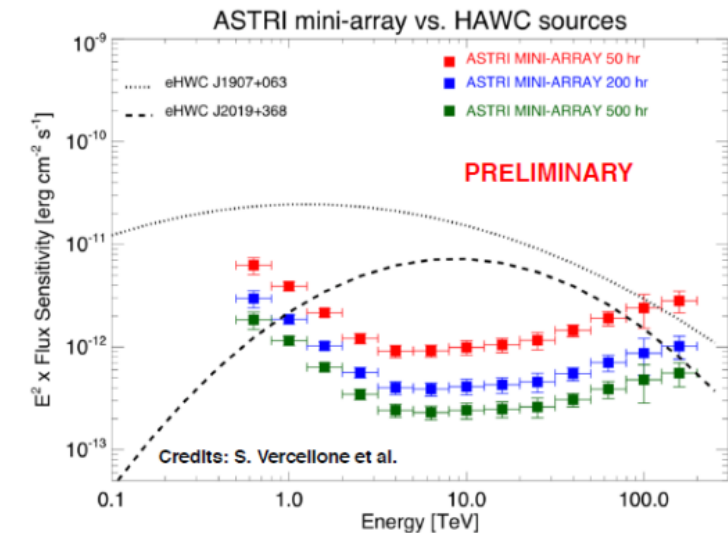
- Restricted number of targets/deep exposures ( $> \sim 200$  h)  $\rightarrow$  strong scientific cases
- Galactic sources: wide FoV  $\rightarrow$  multi-target fields
- Extragalactic sources: survey of a few promising targets at  $> \sim 10$  TeV scale
- Fundamental physics: studies on LIV, EBL, Axion-Like Particles, ...
- Science beyond VHE astronomy also envisaged  $\rightarrow$  Stellar Intensity Interferometry

## Synergies with current VHE Northern Arrays:

- Observations of HAWC sources with much higher angular/energy resolution
- Simultaneous observations with MAGIC and LSTs will be possible

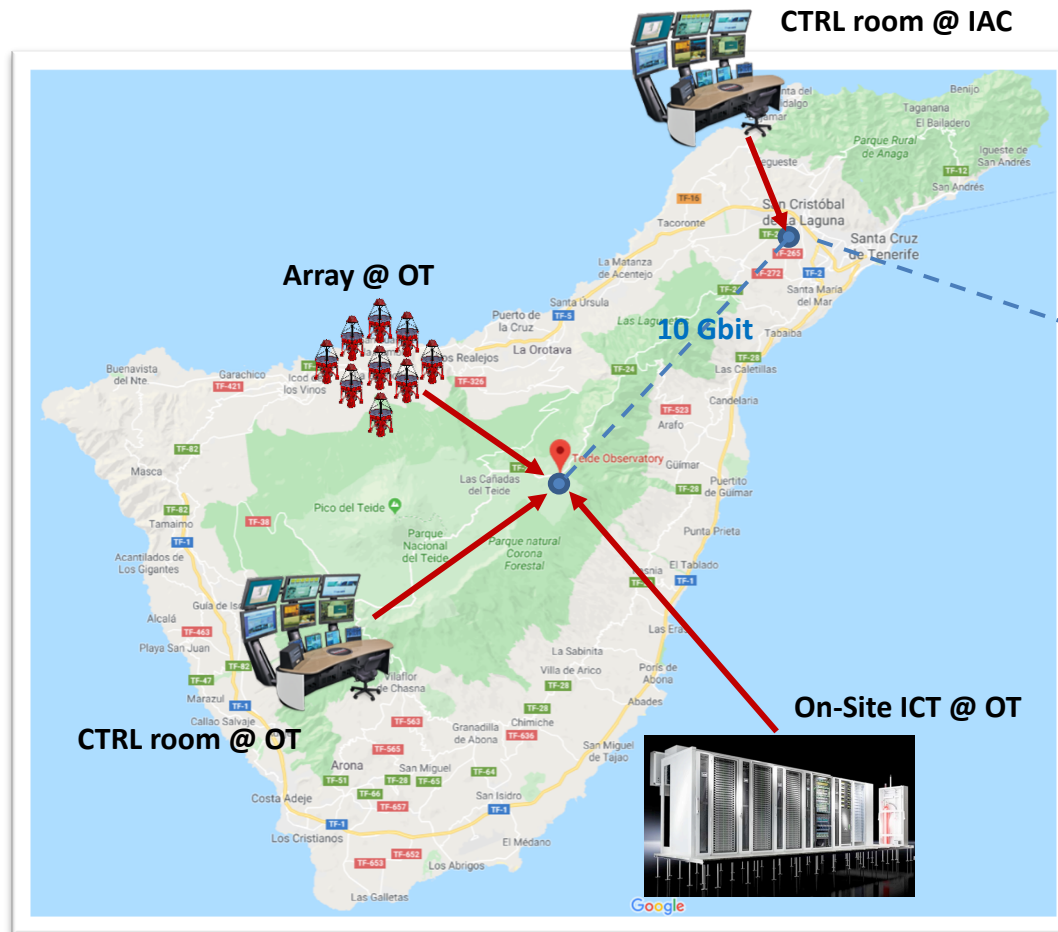


Abeysekara et al. 2019 [arXiv:1909.08609]



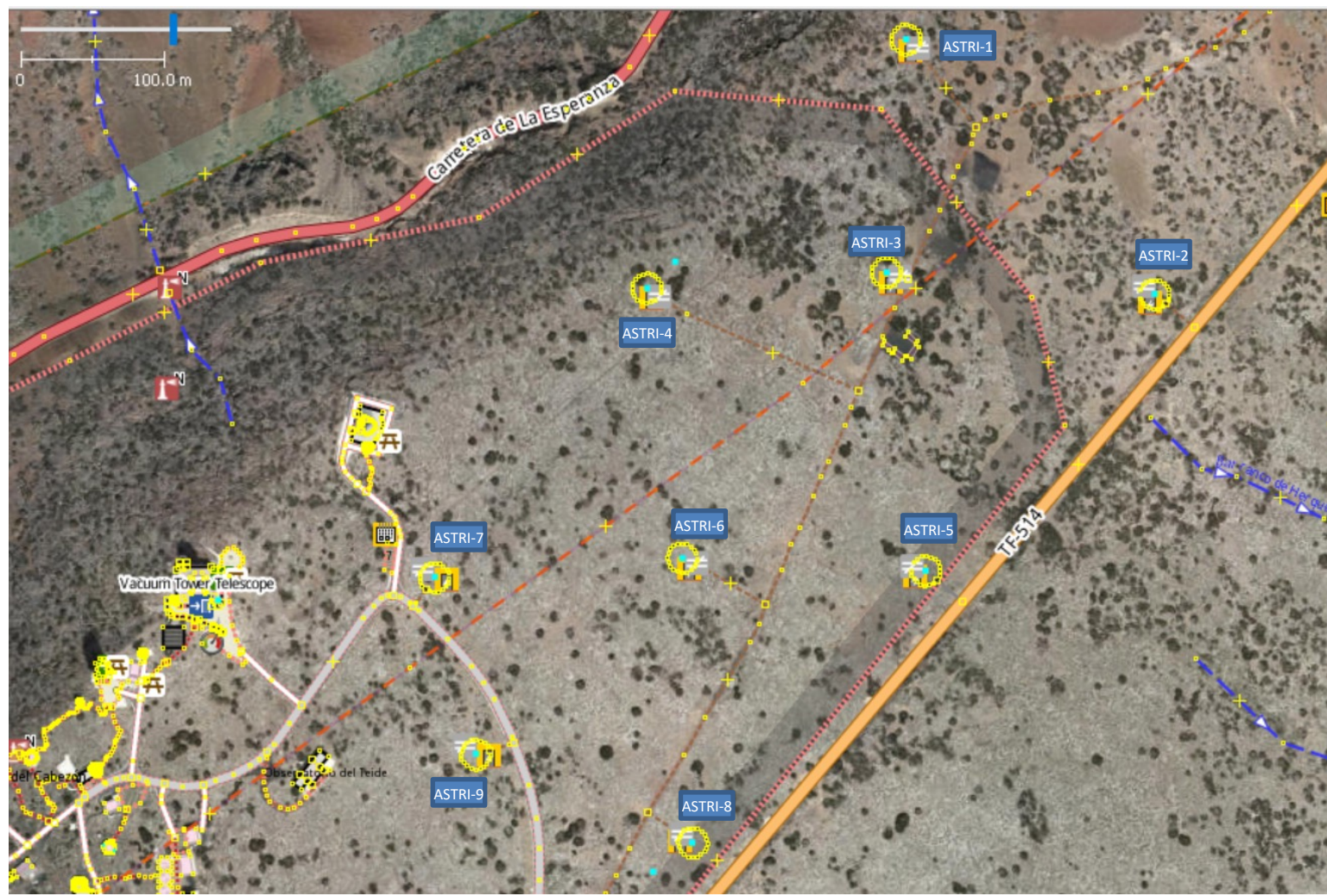
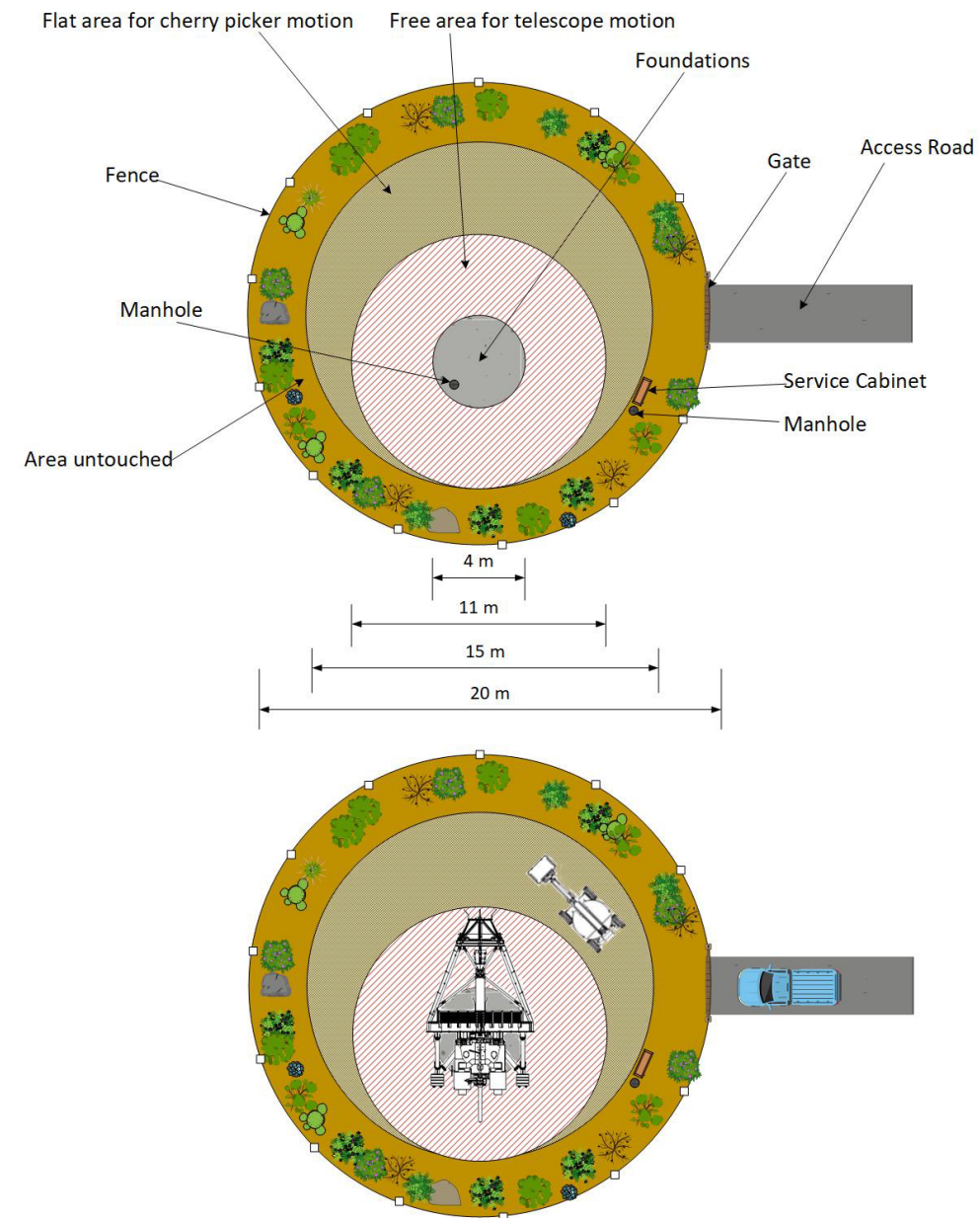


# ASTRI mini-array: The physical layout



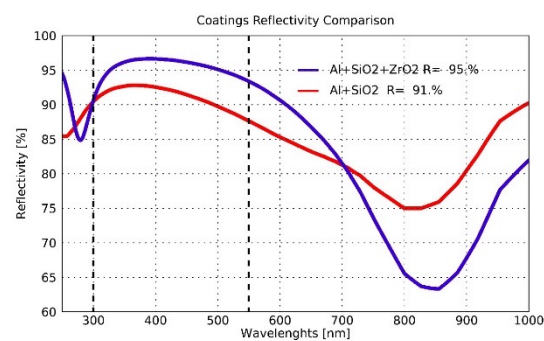
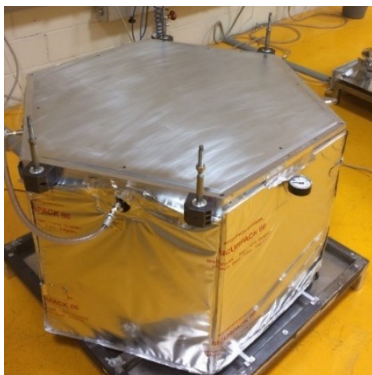
- It will be developed in order to be operated, after the commissioning phase, remotely.
- Data analysis will be performed off-site

## Site ready to host the first telescope in September 2021



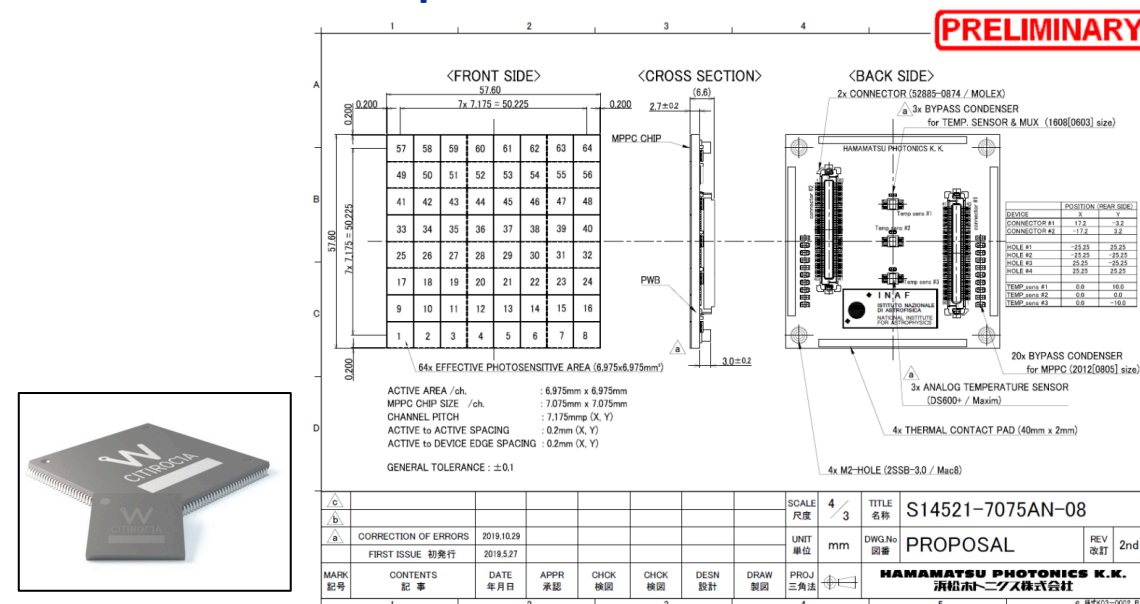
## Mirrors & Camera

M1 and M2 production completed and mirrors packed and ready for shipping



Credits: Media Lario, Flabeg, ZAOT

- Procurement of CITIROC-1A completed
- Procurement of SiPM detectors ongoing
  - Hamamatsu LV3, 7x7 mm, uncoated
- Procurement of Cherenkov Camera
  - tender process started

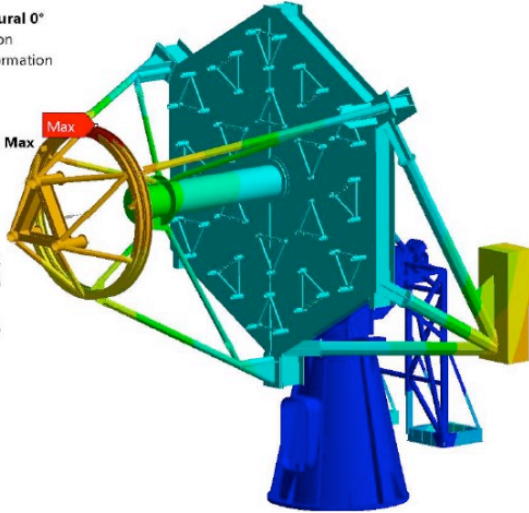


Credits: Hamamatsu Photonics, Weeroc

## Electro-Mechanical Structure

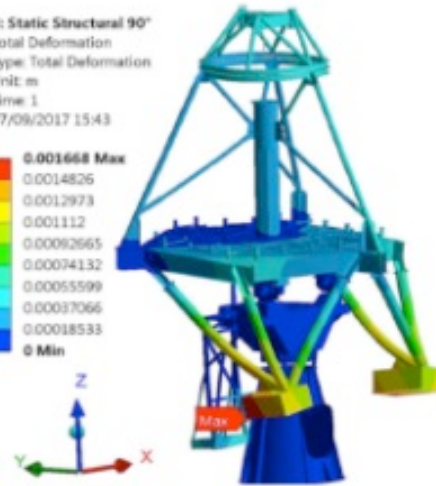
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Total Deformation  
Type: Total Deformation  
Unit: m  
Time: 1

0.0021083 Max  
0.001874  
0.0016398  
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0.00023425  
0 Min

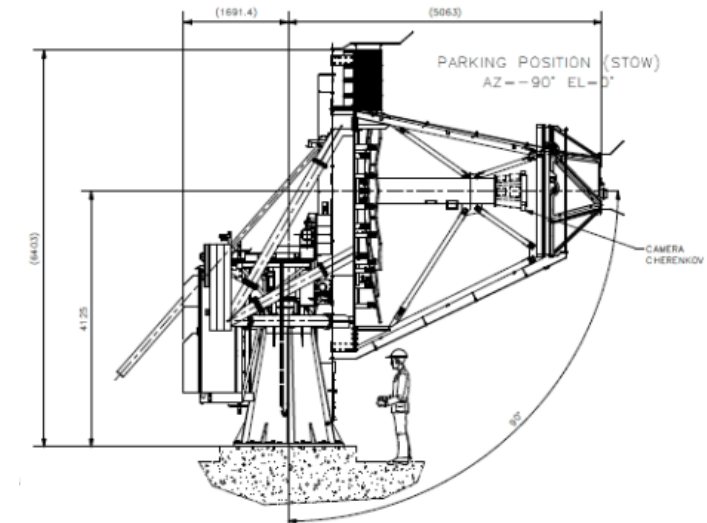
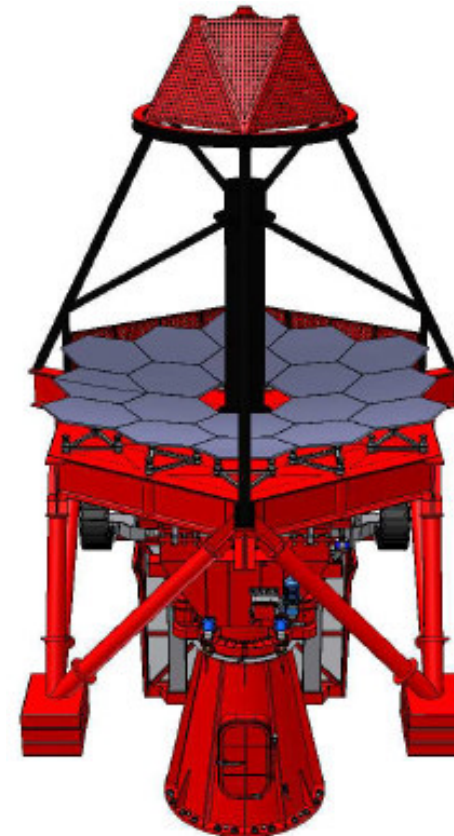


B: Static Structural 90°  
Total Deformation  
Type: Total Deformation  
Unit: m  
Time: 1  
07/09/2017 15:43

0.001668 Max  
0.0014826  
0.0012973  
0.001112  
0.00092665  
0.00074132  
0.00055599  
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0.00018533  
0 Min



- Production of **first three structures COMPLETED**
- Tender for remaining six to be issued



Credits: EIE, Galbiati

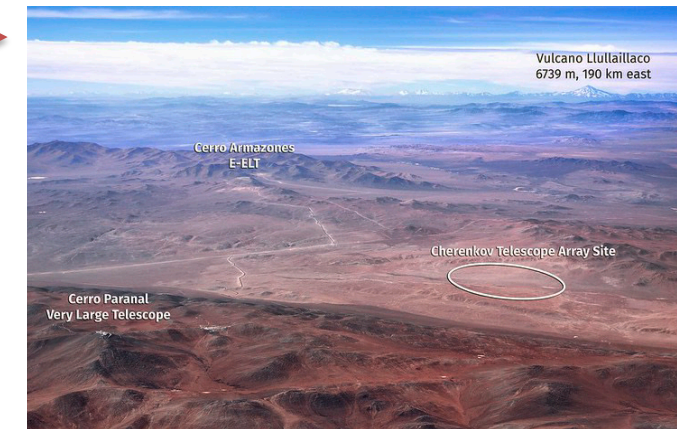
### Design Consolidation

- **Mass reduction (25% → 17 Tons)**
  - Design consolidation of the dish and secondary support to maintain same stiffness lowering telescope weight
  - **Finite Element Analysis** → The structure is able to support seismic loads without suffering damages that will prevent motion.
- **Active Mirror Control (AMC) simplification**
  - No need for AMC during operations
  - AMC radially mounted for easier mirrors integration
  - AMC mounted only in integration phase and for maintenance



## CTA observatory (CTAO):

- Next-generation ground-based Gamma-ray Observatory
- Open observatory designed to operate for 3 decades
- Two sites with total ~100 telescopes (LSTs + MSTs + SSTs)
  - Northern site: La Palma (Canary Islands, Spain)
  - Southern site: Paranal (Chile)
    - 4 Large-size Telescopes (LSTs ~23 meters)
    - 23 Medium-size Telescopes (MSTs ~12 meters)
    - 37-50 Small-size Telescopes (SSTs ~4 meters)
- ~1450 members, ~200 institutes, 31 countries





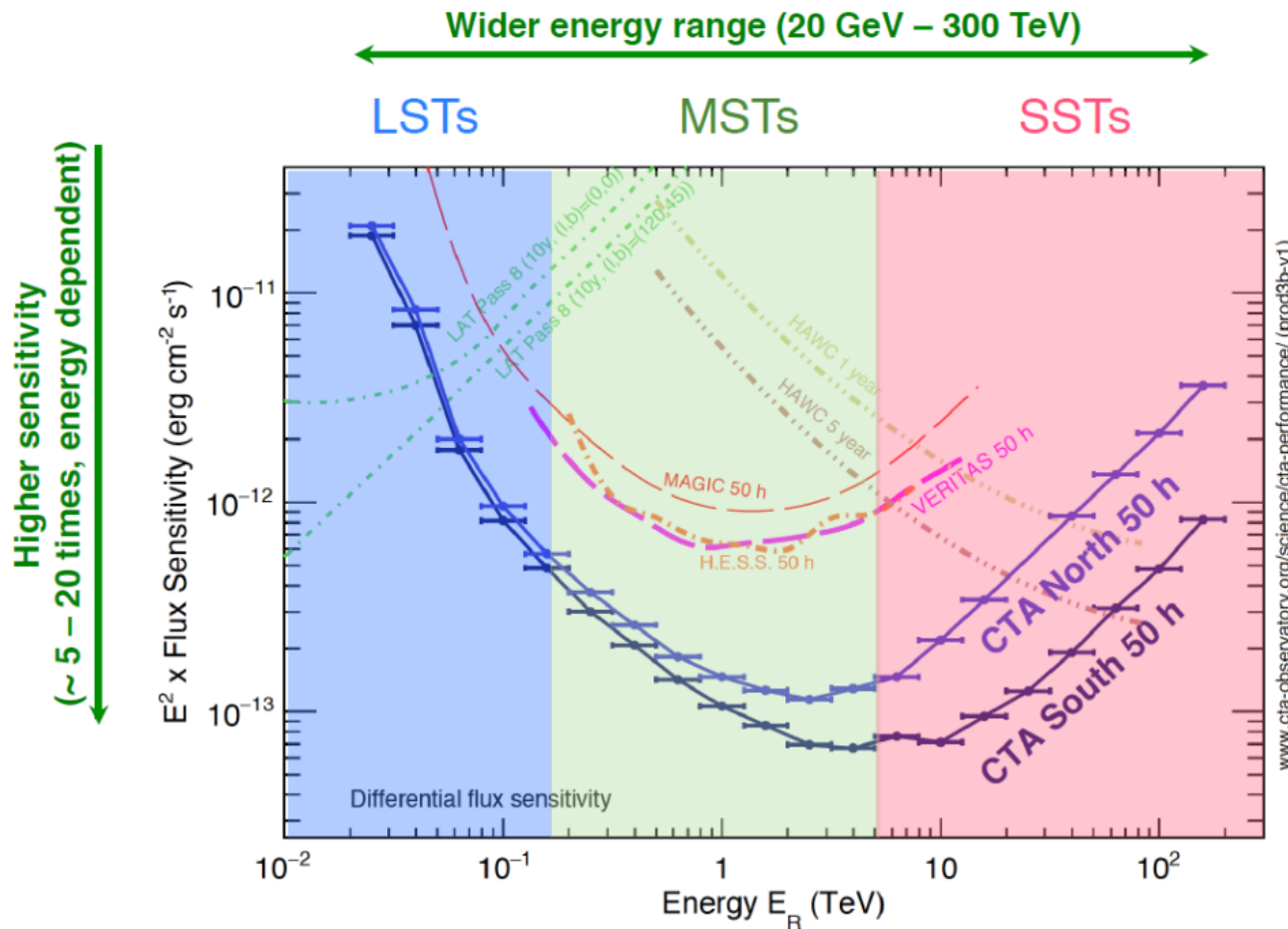


## CTA will address three science themes

1. Cosmic Particle Acceleration
2. Probing Extreme Environments
3. Physics frontiers – beyond the Standard Model

## through several key science programs

- Dark Matter Programme
- Galactic Center Survey
- Galactic Plane Survey
- Large Magellanic Cloud Survey
- Extragalactic Survey
- Transients
- Cosmic-ray PeVatrons
- Star Forming Systems
- Active Galactic Nuclei
- Clusters of Galaxies



www.cta-observatory.org/science/cta-performance/ (prod3b-v1)

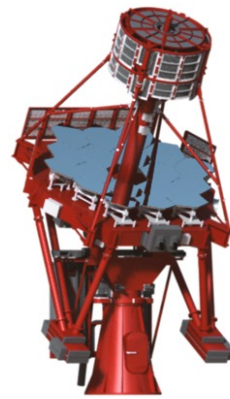
Acharya et al., the CTA Consortium, World Scientific Press (2017) [arXiv:1709.07997]

Science with the Cherenkov Telescope Array. Edited by CTA Consortium <https://www.worldscientific.com/worldscibooks/10.1142/10986>

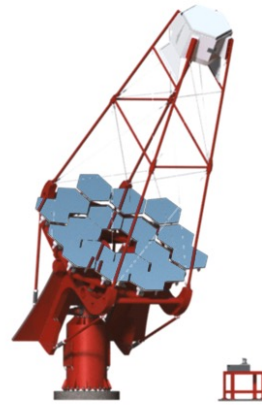
*In the beginning...*



SST-2M GCT



SST-2M ASTRI



SST-1M



**May 2018** → CTAO Council decision to *“to evolve towards a single design for the SST structure and camera”*

**Summer 2018** → Start of the “harmonization” process through a Request for Information.

**March 2019** → Three “proposals” submitted (ASTRI/CHEC, GCT/Obs. de Paris, Consortium 1M) and evaluated by an external committee

**June 2019** → Council resolution on the SST configuration *“..., the CTAO Management proposes that the CTA-SST design be based on the ASTRI-CHEC dual-mirror scheme taking into account the previous experience gained on all designs.”*

**November 2019** → A high level management and implementation plan for the SST telescopes presented at the CTAO Council by ASTRI, CHEC and Obs de Paris teams

# SST Programme Proposal

A high-level implementation plan for the provision of the SSTs

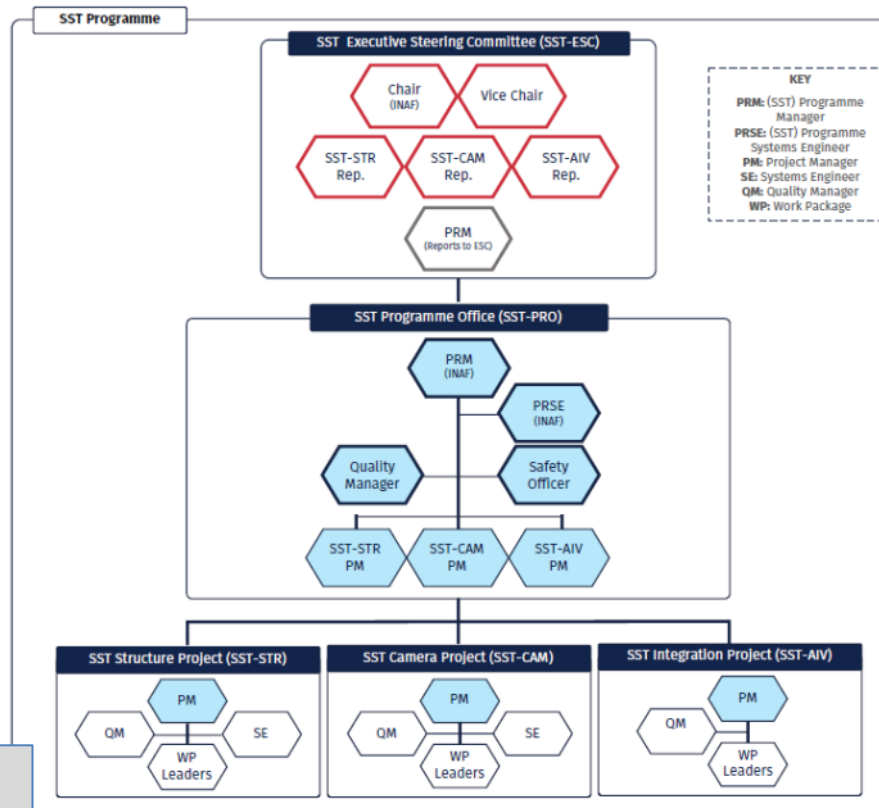
Version  
V2.0 (08-01-2020)

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**OVERALL COST about 50 MEURO**

# High Level Organization of the SST Consortium

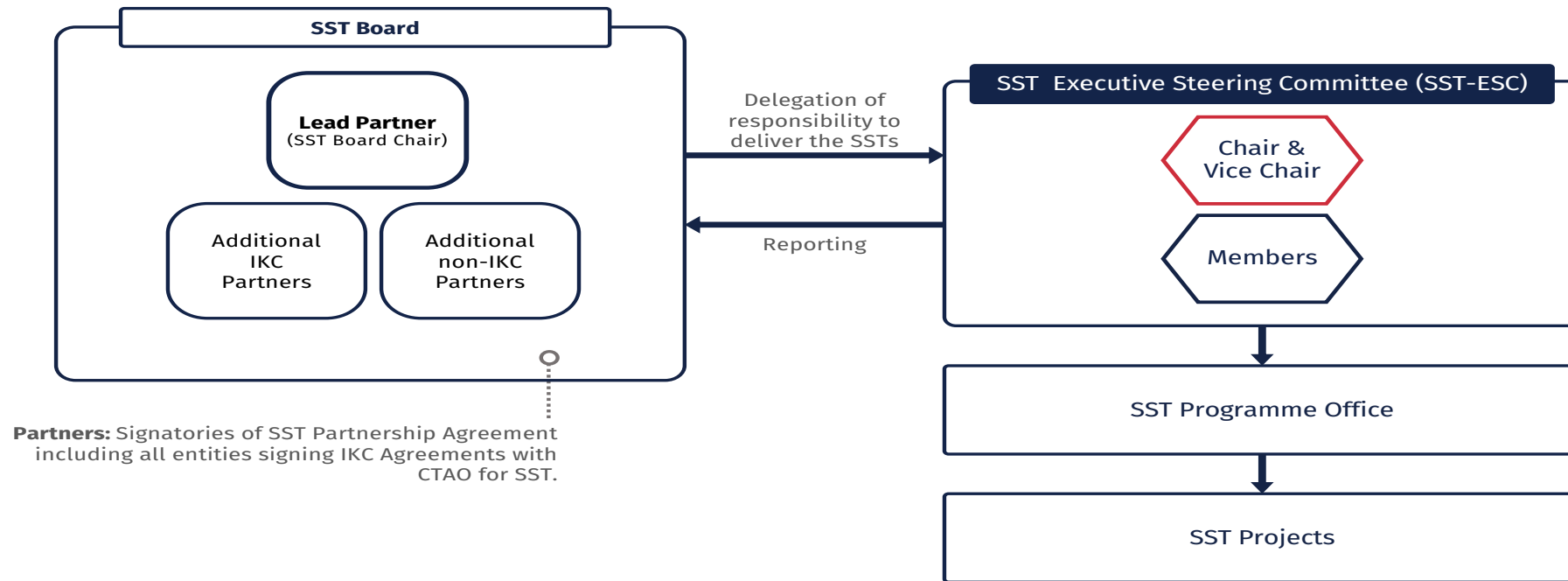


Figure 1: Overview of the SST Governance structure.

**E. M. de Gouveia Dal Pino: member of the SST Executive Steering Committee**

# Interactions with CTAO

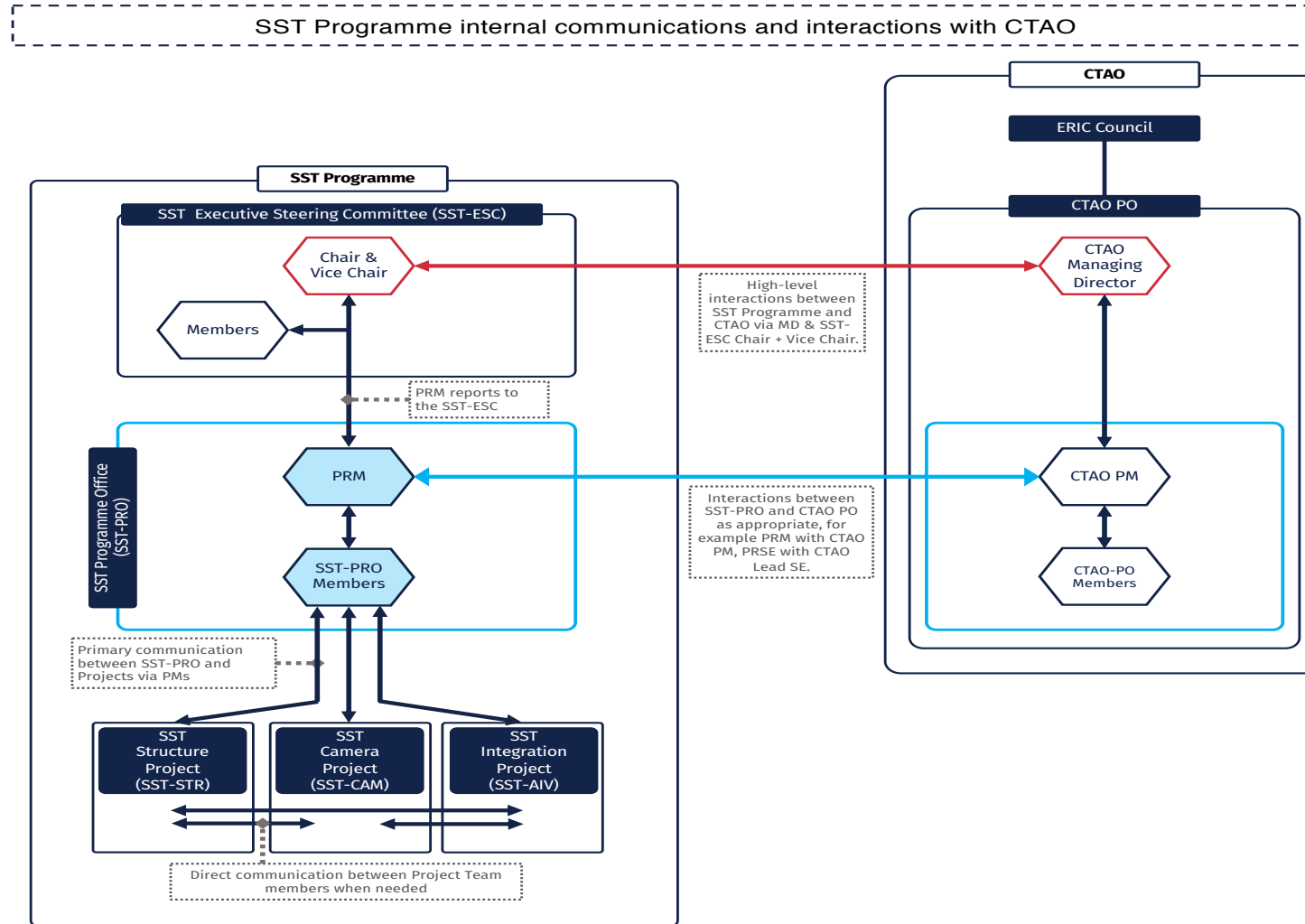
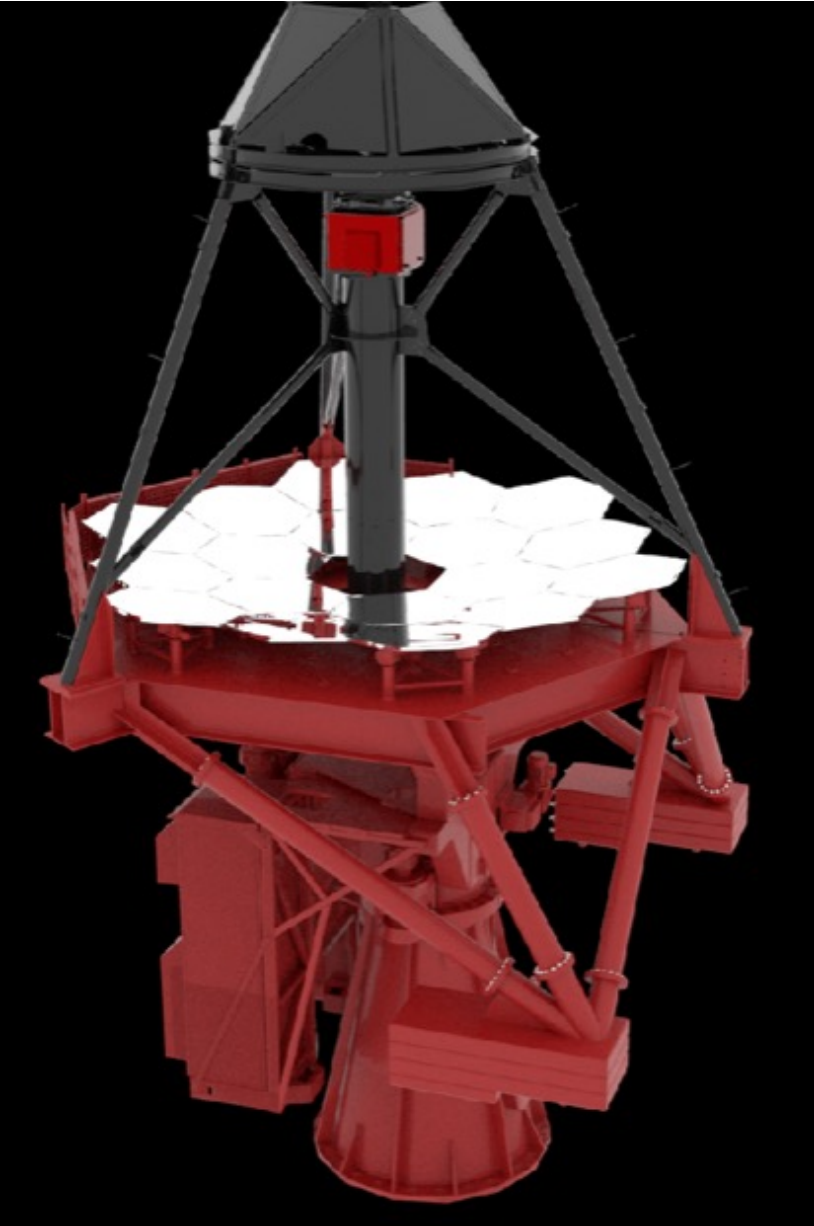


Figure 3: Outline of internal Programme communications and SST Programme interactions with CTAO. Exact scheme to be finalised upon release of CTAO Project Management Plan.



- **Design Review of the ASTRI/CHEC design organized by CTAO with external members (also able to consider different solutions for subsystems)**
- **Optimization of the design for both Structure and Camera and finalization of the design leading to a Critical Design Review and then to production**
- **Preparation of the Implementation and Managements plans → ad interim (waiting for ERIC and IKC) Program Office (and final Steering Committee will be appointed)**
- **Verification of the resources and trade off (in close collaboration with CTAO)**

# BRAZILIANS in the ASTRI & CTA-SSTs *so far*



- **About 29: researchers, post-docs, students, engineers**
- **IAG-USP**
- **EACH-USP**
- **IFUSP**
- **Universidade Mackenzie**
- **UFABC**
- **CBPF**

# BRAZILIAN BUDGET in ASTRI & CTA *so far*



- **FAPESP** : total budget awarded: **R\$ 19.3M**
  - ✓ **Thematic Project 13/10559-5 (Elisabete de Gouveia Dal Pino, IAG-USP): 14,4 M R\$**
    - + **2,7 M R\$ (postdoc scholarships and international advanced school)**



# OUR PARTICIPATION *history so far*



- **2016 ASTRI MOU** involving USP, FAPESP and INAF: roles and responsibilities of the parties transfer of technology, intellectual property use, governance and maintenance of the MINI-ARRAY
- **1.6 MEuro** issued to INAF for 3 ASTRI Structures (2017)
- **Several scholarships** (in our Project: 6 postdocs, 2 JPs, 6 PhDs, 1 M, 5 ICs)
- **Training of Human Resources: SPSAS-HighAstro CTA - Sao Paulo Advanced Science School on High Energy and Plasma Astrophysics at the CTA Era (2017)**
- **Development of the Science Working Packages of CTA & ASTRI Mini-Array**
- **Editorial board** of ASTRI papers (ASPO)
- **Engineers Renato Gimenes and Carlos Fermino: Instrumentation development**

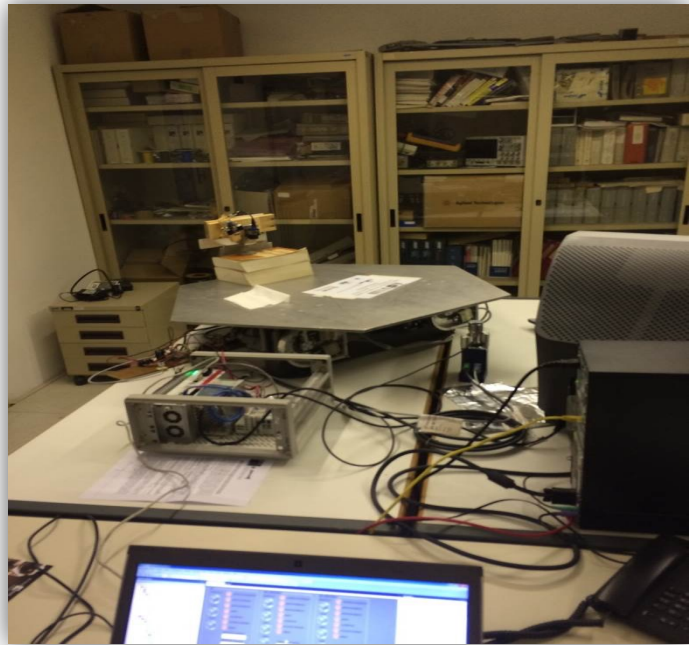
# BRAZILIAN ENGINEERS @ ASTRI: structure/cam/software development

## Active optics tests (Torino)

- With Daniele Gardiol & Federico Russo

## Performed task:

- Segment motion
- Data collection and calibration
- Repeatability verification

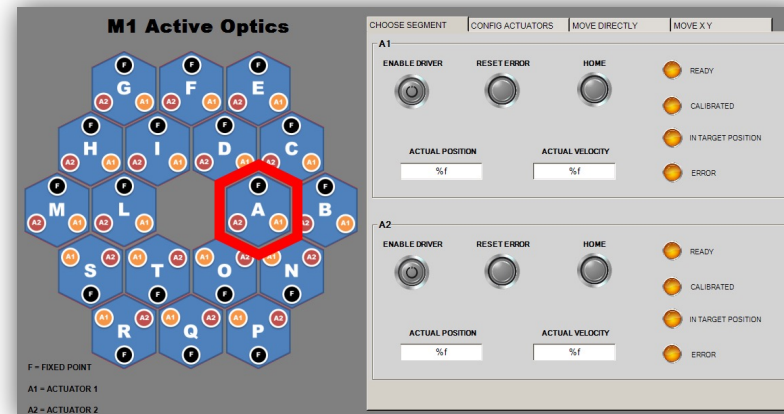


## M1 Interface elaboration (Catania)

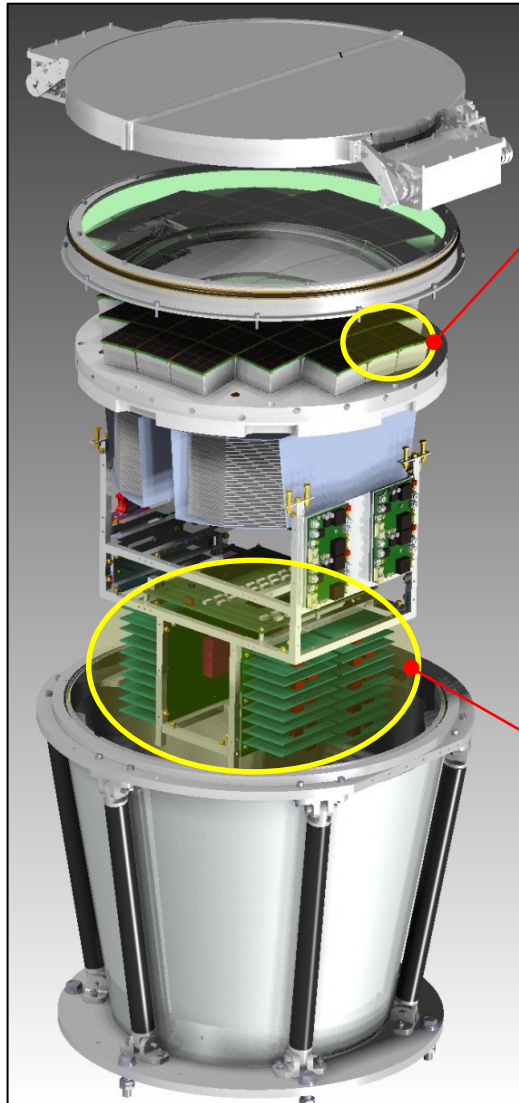
- W/ Matteo Munari, Salvo Scuderi, Enrico Giro, Luca Stringhetti, Elisa Antolini

## Performed task:

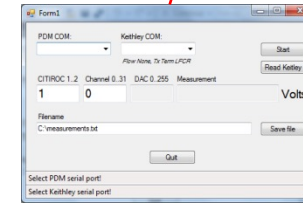
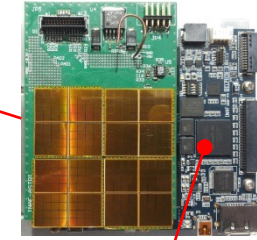
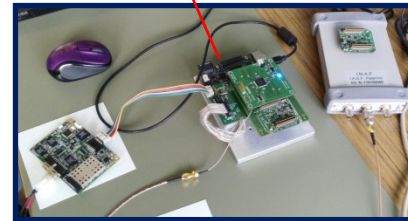
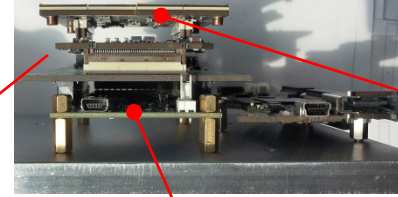
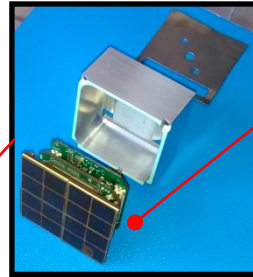
- Actuator configuration as in CANopen
- Screens preparation for driving motors
- interface implementation with users, scheduling tests, etc.
- Teaching on operation of logics for future change and implementation
- Teaching hardware & software Beckhoff



# BRAZILIAN ENGINEERS @ IASF Palermo

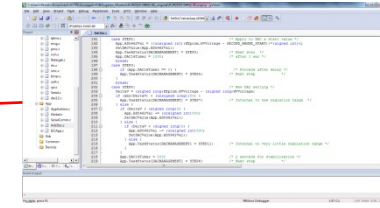
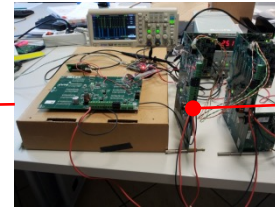


**ASTRI Camera**



## **PDMs – Photodetector Modules**

- *Hardware Improvements and EMI Countermeasures*
- *Calibrations, Test Set-Ups and PC Applications*



## **VDB – Voltage Distribution Box**

- *Hardware and Firmware Corrections and Improvements*
- *Test Set-Ups, Test Jigs and PC Control Software*
- *Calibration Routines and Climate Tests*
- *DC/DC Converter Design for New SiPMs Biasing*

# NEXT STEP PARTICIPATION in SSTs for



- **Brasil in the SST Program:** ASTRI+CHEC (editorial board document)
- **Currently: ~60 members in CTA-Br ( ~29 -> in this Project from IAG-USP, EACH-USP, Mackenzie, UFABC, IFUSP, CBPF)**
- **Plans for manufacturing in Brazil at least 3 of the 37 ASTRI SSTs structures of CTA-South array**
- **Approximate total cost: ~12 MReais**
- **Technical Advisory: Eng. Carlos Fermino (eFe – Tecnologia**  
**(<http://www.incubadora-araraquara.com.br/empresas-incubadas/residentes>)**
- **Potential Companies: ZANINI; IESA; CITROTEC;\_SERMATEC**



Universidade de São Paulo



# CTA Precursor towards CTA



Thank You!